

# Human and Robotic Exploration Activities at Small Bodies: Opportunities for Science

Paul A. Abell Astromaterials Research & Exploration Science Directorate NASA Johnson Space Center Houston, Texas

Stakeholder Engagement on the Global Exploration Roadmap: Focus on Science NASA Ames Research Center Moffett Field, CA
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## Outline

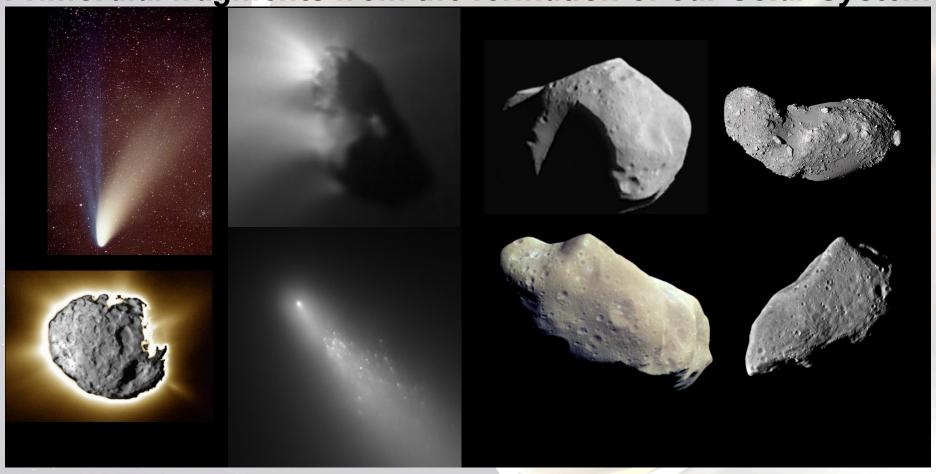


- Introduction to Small Bodies
- Small Body Focal Areas
- Planetary Decadal Science Themes
- Role of Robotic Precursors
- Science Enabled by Human Presence
- > Future Considerations

## Small Bodies = Comets and Asteroids 🍇



Primordial fragments from the formation of our Solar System



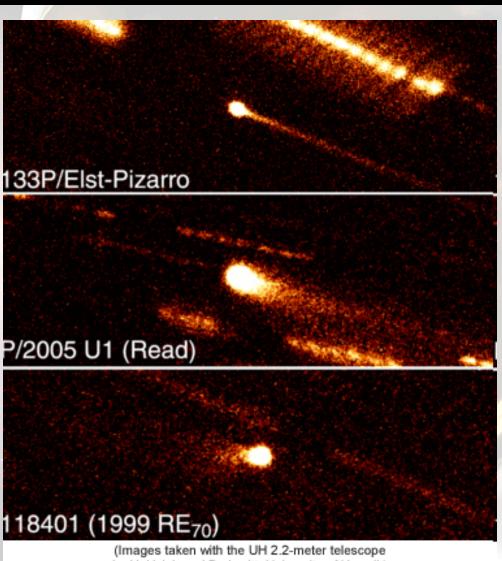
Comets contain volatiles in the form of ices and can produce visible atmospheres (coma)

Asteroids generally lack active ices and appear to be essentially inert

#### **Examples of Dormant Comet/Active Asteroids**



- Recent discoveries of main belt comets
- Small bodies are an extremely diverse population of objects
- Represent several avenues for scientific research



## Detection of Water Vapour at Asteroid (1) Ceres



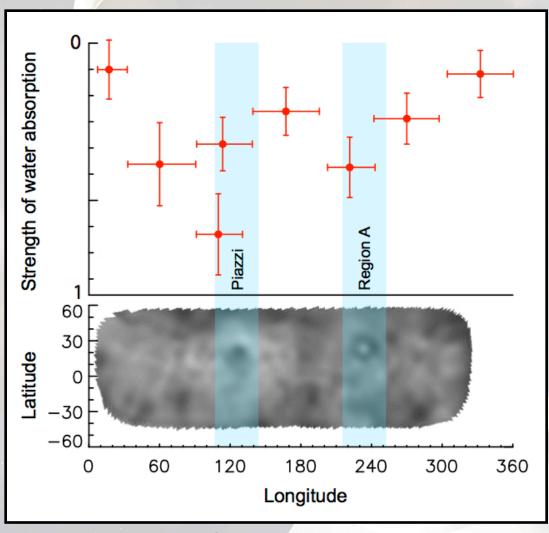
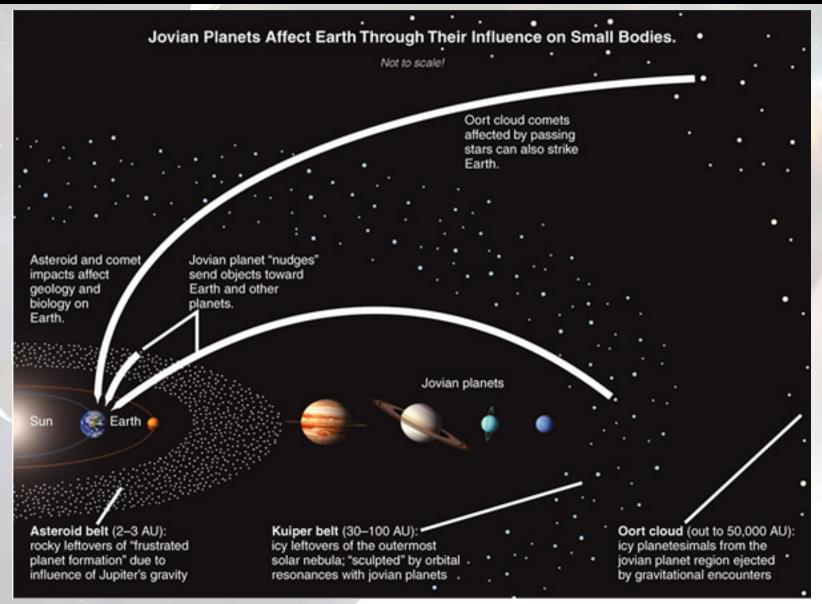


Image credit from Küppers et al., 2014

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## **Small Body Source Regions**



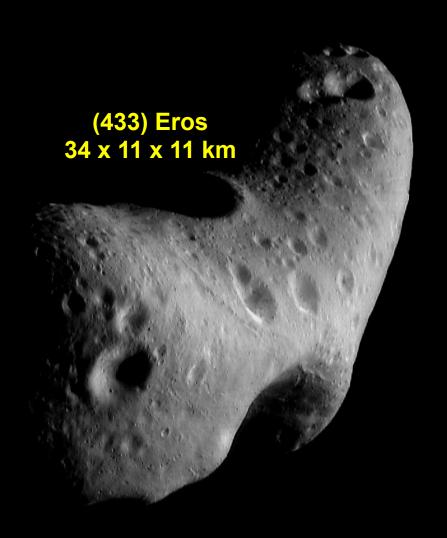


#### **Potential Small Body Destinations**



- Near-Earth Objects (NEOs) and Phobos/Deimos are small airless bodies with "similar" physical characteristics
- However they represent distinct and separate destinations for robotic and human exploration
  - NEOs any small body passing within 1.3 Astronomical Units (AU) of the Sun
    - Note that 90 95% of the NEO population are Near-Earth Asteroids (NEAs)
    - 5 10% of NEOs may be extinct or dormant comets
       (i.e., no evidence of activity)
    - NEAs are the targets for Human Exploration
  - Phobos and Deimos are natural satellites of Mars at ~1.52 AU

#### **NEAs to scale with the Martian moons**





Phobos 26 x 22 x 18 km

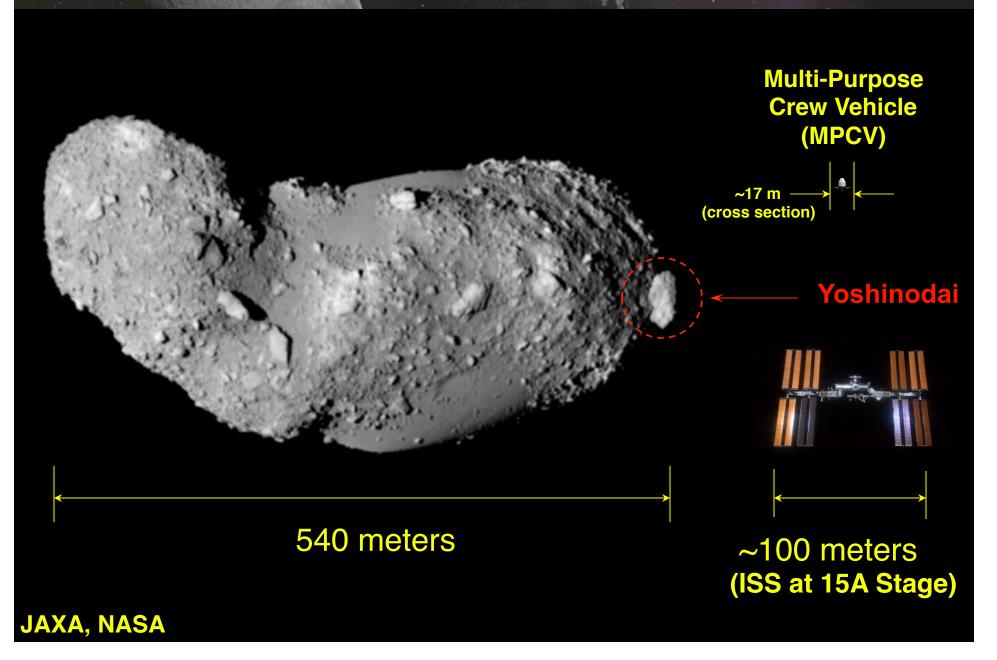


**Deimos 15 x 12 x 10 km** 

(25143) Itokawa 0.54 x 0.29 x .20 km

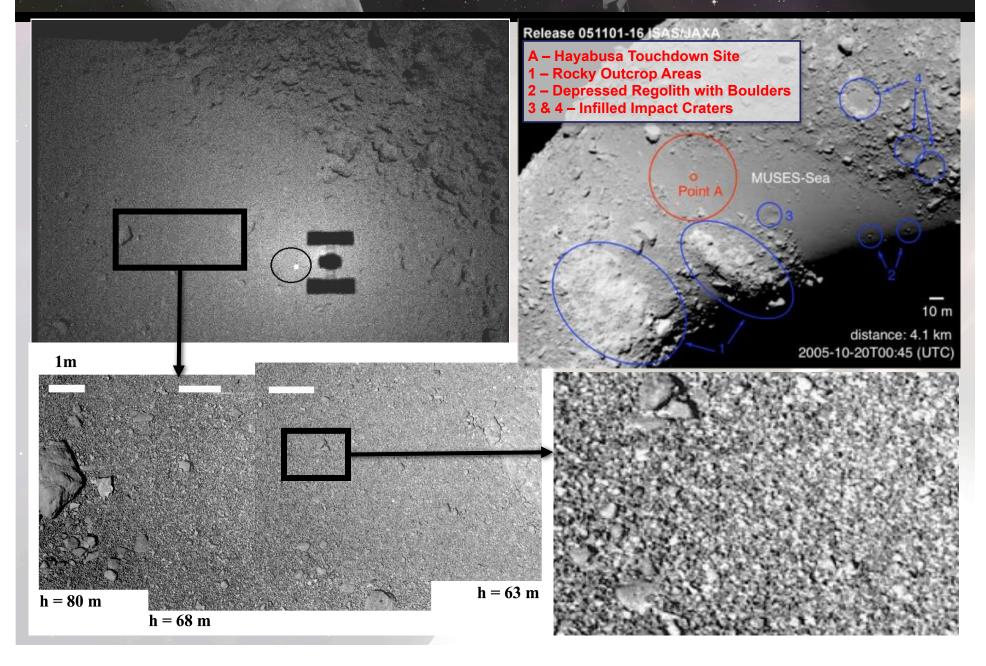
## Asteroid Itokawa, ISS, and the MPCV





#### Hayabusa Touchdown Site at Itokawa





# **Small Body Focal Areas** - Exploration - Science - Resources - Planetary Defense

## Overlap of Small Body Focal Areas



#### **Human Operations**

Internal Structure (monolith vs. rubble pile)
Sub-surface properties
General mineral, chemical composition

#### Science

Internal Structure (monolith vs. rubble pile)
Sub-surface properties
Detailed mineral, chemical, isotopic composition

#### Intersection of all

Location (position prediction/orbit)
Size (existence of binary/ternary)
Rotation rate & pole orientation
Particulate environment/Debris field
Electrostatic charging & plasma field
Thermal environment
Gravitational field structure
Mass/Density estimates
Surface morphology and properties
Regolith mechanical & geotechnical properties

Internal Structure (monolith vs. rubble pile)
Sub-surface properties (-> Beta)
General mineral, chemical composition

Detailed mineral, chemical composition

**Planetary Defense** 

**Resource Utilization** 

## Planetary Decadal Science Themes



- Building New Worlds -- understanding solar system beginnings
  - What were the initial stages, conditions and processes of solar system formation and the nature of the interstellar matter that was incorporated?
  - What governed the accretion, supply of water, chemistry, and internal differentiation of the inner planets and the evolution of their atmospheres, and what roles did bombardment by large projectiles play?
- Planetary Habitats -- searching for the requirements for life
  - What were the primordial sources of organic matter, and where does organic synthesis continue today?
- Workings of Solar Systems -- revealing planetary processes through time
  - What solar system bodies endanger and what mechanisms shield Earth's biosphere?
  - How have the myriad chemical and physical processes that shaped the solar system operated, interacted, and evolved over time?

## **Small Body Science**



- Small bodies are the left over primitive materials from the earliest stages of Solar System formation
  - NEAs sample a wide diversity of material types and reservoirs
  - Context of Small Body origins and evolution
- In situ science enables a better understanding of these bodies origin/dynamical history and early Solar System formation
- Proximity science investigations ("orbit" or station keeping modes)
- Surface and sub-surface science investigations (regolith/material properties drivers) to identify
  origins, investigate impact history, understand space weathering effects, etc.
- > Sample return from these bodies informs the nature of the material composition, thermal, oxidation state, and collisional histories
  - Identification and characterization of samples
  - Acquisition under low-gravity regimes with regolith and material property drivers
  - Storage and containment of samples (i.e., cold/frozen samples)
- Identification of volatile and organic materials has major implications for astrobiology
- Delivery of pre-biotic molecules and volatiles to Earth via NEAs (e.g., origin of Earth's H<sub>2</sub>O)

#### Small Body In Situ Characterization



- Prior to sending a piloted mission to a NEA, additional in situ characterization of the target is required
  - Obtain basic reconnaissance to assess potential hazards that may pose a risk to both vehicle and crew (e.g., Ranger, Lunar Orbiter, Surveyor, LRO have done for the Moon)
  - Some NEAs may have physical characteristics that would make them unsuitable as targets for early human exploration
  - Fill in Strategic Knowledge Gaps (SKGs)
- Precursor missions would also assess the NEA for future activities to be conducted by the crew and their assets to maximize mission efficiency
  - Proximity operations (spacecraft, robotic assets, crew on EVA)
  - Surface activities (scientific and engineering investigations)
  - Touchdown site selection
  - Sample identification and collection

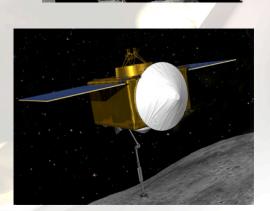
#### Small Body In Situ Characterization



- Build on lessons learned from other small body missions
- Strong overlap with Science, Exploration, Resource Utilization, and Planetary Defence
- Past, current, and future mission examples
  - NEAR Shoemaker, Hayabusa,
  - Rosetta, Hayabusa 2, OSIRIS REx,
  - AIDA, ISIS







#### **Robotic Precursor Activities**



- The objectives for the robotic mission would be oriented towards data collection to prepare for Human exploration
  - High-resolution gravity field, density measurements (rubble pile vs. monolith), and topographic mapping
  - Surface composition mapping, resource characterization and distribution
  - Dynamical environment (radiation, dust levitation, plasma fields interactions, etc)
- The precursor should have some level of direct interaction with the surface (e.g., landers, hoppers, penetrators, etc.)
  - Near-surface regolith geotechnical properties (composition/size distribution, stability, porosity, compaction, etc.)
- Robotic sample return prior to the Human mission is not necessary as a risk- or cost-reduction activity
  - Dust mitigation and contamination protocols can be based on previous small body experience gained from the Hayabusa, Rosetta, Hayabusa 2, and OSIRIS REx missions
  - The Human crew is more suited to obtain macroscopic and diversified samples

## **Robotic Precursor Support**



## The precursor may also be extremely valuable during (and after) human mission operations

- Experience has shown that small body robotic probes can be operational for 10+ years (e.g., Stardust, Deep Impact, etc.)
- The precursor can provide situational awareness (bird's eye view) and act as a data/communication relay for any assets out the line of sight

 It can also act as a data rely for equipment/instruments left behind on the surface for long duration investigations (e.g., geophysical stations, ISRU

operations, etc.)



## Science Enabled by Human Presence



#### > Humans are able to perform complex tasks and make rapid decisions in real time

- Ideally suited to performing detailed and complex tasks which will be required in low-gravity environmental conditions at planetary surfaces
- Enables a more efficient and higher quality science return
- Recognizing and collecting scientifically valuable samples for return to Earth is one of the most important tasks at a small body (e.g., age dating, evolution, parent body processes, etc.)

#### > The support infrastructure will further enhance science return

- Combination of teleoperated and autonomous systems furthers the capabilities of human interaction and investigation at the target
  - Independent and simultaneous operations
  - Detailed coordinated activities
- Computer capabilities and power of spacecraft systems can be utilized on station
  - Radar tomography of small body interiors
  - High bandwidth for data transfer and communication
- Enhanced cargo capacity for operations and samples
  - Delivery of science and engineering equipment to be used during EVAs
  - Capability for larger return and diversity of samples
  - Options for containment (e.g., cold temperatures)



Source: NASA/AMA, Inc.

#### Science Investigations



#### Determine the nature of the small body surface geology and mineralogy

- This provides context and sample selection information for the returned samples for which these
  measurements will be conducted in Earth laboratories
- Information on bulk composition and origin of materials (meteorite affinities/parent bodies)
- Recognition of any alteration processes (fresh vs. weathered)
- Can be done via hand held remote sensing or in situ instruments

#### Characterize the regolith of the small body at local and global scales

- Interpret the processes that have formed and modified it
- Study the nature, structure, and degree of maturation
- Particle size frequency distributions needed to characterize microgravity geology
- Understand the geotechnical properties (cohesion, density, porosity, etc.)
- Characterize the thermal inertia and mobility response
- Best suited to geophysical equipment, subsurface excavation (trenching/drilling/core), probes, etc.

#### Identify and characterize the presence and distribution of volatiles or organics

- Investigate at surface, near-surface, and at depth
- Can be done with hand held remote sensing or in situ instruments
- Detailed investigation on returned samples relevant for astrobiology

#### Science Investigations



- Determine the near surface and interior structure at regional and global scales
  - Identification of rubble pile or monolithic structures. Also determination of interior component size frequency distribution (large pieces vs. meter scale bits)
  - Seismic surveys and radar tomography
  - Ground penetrating radar
- Characterize the dust, plasma, and electrostatics of the airless body
  - Presence of any exosphere and potential interaction during solar activity
  - Characterize micrometeorite flux (source of dust and any sinks)
  - Combination of in situ probes and spacecraft sensors
- Measure and monitor radiation environment on site and characterize the response of the small body surface material response
  - Identify shielding opportunities of small body materials
  - Characterize any secondary radiation emitted from the small body
  - Rad sensors and tissue equivalent detectors

#### **Future Considerations**



- Opportunities exist for enhanced science capability with the systems and elements needed for missions to the Moon, Mars, and small bodies
  - Multi-purpose Crew Vehicle (MPCV)
  - Space Launch System (SLS)
  - Exploration Augmentation Module (EAM)
  - Deep Space Habitat (DSH)
  - Multi-Mission Space Exploration Vehicle (MMSEV) (i.e., excursion vehicle)
- Foster international, commercial, and academic cooperation for both robotic precursor and human missions
  - JAXA, ESA, Roscomos, CSA, and other exploration agencies
  - Space X, Planetary Resources, Deep Space Industries, Bigelow Aerospace, etc.
  - Scientific and Engineering communities



